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**Abstract:** The aim of this paper is to characterize a Riemannian 3-manifold  $M^3$  equipped with a semisymmetric metric  $\xi$ -connection  $\tilde{\nabla}$  with  $\rho$ -Einstein and gradient  $\rho$ -Einstein solitons. The existence of a gradient  $\rho$ -Einstein soliton in an  $M^3$  admitting  $\tilde{\nabla}$  is ensured by constructing a non-trivial example, and hence some of our results are verified. By using standard tensorial technique, we prove that the scalar curvature of  $(M^3, \tilde{\nabla})$  satisfies the Poisson equation  $\Delta R = \frac{4(2-\sigma-6\rho)}{\rho}$ .

Keywords: Riemannian manifolds;  $\rho$ -Einstein solitons; Einstein manifolds; Poisson equation

MSC: 53E20; 53C25; 53C21

### 1. Introduction

The Ricci and other geometric flows are active topics of current research in mathematics, physics and engineering. The Ricci flow [1] is defined on a Riemannian *n*-manifold  $(M^n, g)$  by an evolution equation for metric g(t) of the form  $\frac{\partial g}{\partial t} = -2S$ , where S is the Ricci tensor of  $M^n$  and t indicates the time. The metric g on  $M^n$  satisfies the Ricci soliton (in short, RS) equation  $\mathcal{L}_E g + 2\sigma g + 2S = 0$ , where *E* is a vector field on  $M^n$ ,  $\sigma \in \mathcal{R}$  (the set of real numbers), and  $\mathcal{L}_E$  represents the Lie derivative operator in the direction of E on  $M^n$ . A RS is called expanding (steady or shrinking) if  $\sigma > 0$  ( $\sigma = 0$  or  $\sigma < 0$ ). If E = 0or Killing, then the RS is called a trivial RS, and  $M^n$  becomes an Einstein manifold. Thus the RS is a basic generalization of an Einstein manifold [2]. If  $\mathcal{F}$  is a smooth function such that  $E = \mathcal{DF}$  for the gradient operator  $\mathcal{D}$  of g, then the RS is described as a gradient Ricci soliton (GRS), E is referred to as the potential vector field, and  $\mathcal{F}$  is called the potential function. Thus, the RS equation becomes  $Hess\mathcal{F} + \sigma g + S = 0$ , where  $Hess\mathcal{F}$  is the Hessian of  $\mathcal{F}$  and  $(Hess\mathcal{F})(\zeta_1,\zeta_2) = g(\nabla_{\zeta_1}D\mathcal{F},\zeta_2)$  for all vector fields  $\zeta_1$  and  $\zeta_2$  on  $M^n$ . Here,  $\nabla$ stands for the Levi-Civita connection.

The notion of Ricci–Bourguignon flow, a natural generalization of Ricci flow, has been proposed in [3] and is described on an  $M^n$  as:

$$\frac{\partial g}{\partial t} = -2(S - \rho Rg), \quad g(0) = g_0, \tag{1}$$

where *R* is the scalar curvature and  $\rho \in \mathcal{R}$ . It is to be noticed that for the specific values of  $\rho$ , the following cases for the tensor  $S - \rho Rg$  appeared in (1) [4] are obtained:

- $\rho = \frac{1}{2}$ , the Einstein tensor  $S \frac{R}{2}g$ , (for Einstein soliton), (i)

  - (ii)  $\rho = \frac{1}{n}$ , the trace-less Ricci tensor  $S \frac{R}{n}g$ , (iii)  $\rho = \frac{1}{2(n-1)}$ , the Schouten tensor  $S \frac{R}{2(n-1)}g$ , (for Schouten soliton),
  - (iv)  $\rho = 0$ , the Ricci tensor *S* (for RS).



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An 
$$(M^n, g)$$
,  $n \ge 3$  is said to be a  $\rho$ -Einstein soliton (or  $\rho$ -ES)  $(g, E, \sigma, \rho)$  if

$$\pounds_E g + 2S + 2(\sigma - \rho R)g = 0. \tag{2}$$

Similar to the RS, a  $\rho$ -ES is called expanding (steady or shrinking) if  $\sigma > 0$  ( $\sigma = 0$  or  $\sigma < 0$ ). If E = DF, then  $(M^n, g)$  is called a gradient  $\rho$ -Einstein soliton (or gradient  $\rho$ -ES). Hence, (2) takes the form

$$Hess\mathcal{F} + S + (\sigma - \rho R)g = 0, \tag{3}$$

where  $Hess\mathcal{F}$  denotes the Hessian of  $\mathcal{F} \in C^{\infty}(M^n)$  and defined by  $Hess\mathcal{F} = \nabla \nabla \mathcal{F}$ . Recently,  $\rho$ -Einstein solitons have been studied by several authors, such as [5–12]. On the other hand, we recommend the papers [13–19] for the studies of Ricci, Yamabe, Ricci-Yamabe,  $\eta$ -Ricci-Yamabe and quasi-Yamabe solitons on different geometric structures.

In this paper, we have made an effort to the solitonic study of a 3-dimensional Riemannian manifold  $M^3$  equipped with a semi-symmetric metric  $\xi$ -connection  $\tilde{\nabla}$ . To achieve the goal, we present our work as follows: In Section 2, we gather the basic information of a Riemannian 3-manifold equipped with a semi-symmetric metric  $\xi$ -connection  $(M^3, \tilde{\nabla}, g)$ , definitions and Lemmas. The properties of  $\rho$ -ES in  $(M^3, \tilde{\nabla}, g)$  are studied in Section 3. We address the properties of gradient  $\rho$ -ES in  $(M^3, \tilde{\nabla}, g)$  in Section 4. In the last section, we model a non-trivial example of  $(M^3, \tilde{\nabla}, g)$  admitting a gradient  $\rho$ -ES, and prove our results.

# 2. Riemannian Manifolds with a Semi-Symmetric Metric $\xi$ -Connection

In 1970, Yano [20] investigated the properties of a semi-symmetric metric connection  $\tilde{\nabla}$  on Riemannian *n*-manifolds  $M^n$  and defined by  $\tilde{\nabla}_{\zeta_1}\zeta_2 = \nabla_{\zeta_1}\zeta_2 + \eta(\zeta_2)\zeta_1 - g(\zeta_1,\zeta_2)\xi$  for all  $\zeta_1$  and  $\zeta_2$  on  $M^n$ , where  $\eta$  is a 1-form associated with the unit vector field  $\xi$  such that  $g(\xi,\xi) = \eta(\xi) = 1$  and  $g(\zeta_1,\xi) = \eta(\zeta_1)$ . Later, the properties of the semi-symmetric metric connection  $\tilde{\nabla}$  have been explored by several researchers. One of these properties is the curvature invariant respecting to the semi-symmetric metric connection  $\tilde{\nabla}$  and the Levi–Civita connection  $\nabla$ . For example, the conformal curvature tensors corresponding to the semi-symmetric connection (Yano's sense) and the Levi–Civita connection coincide. Similar results for different curvature tensors have been established by many geometers. A connection  $\tilde{\nabla}$  is said to be semi-symmetric metric  $\xi$ -connection if and only if  $\tilde{\nabla}\xi = 0$ . Afterwards, the properties of semi-symmetric metric  $\xi$ -connection have been studied in [21–24].

In an  $(M^n, \tilde{\nabla}, g)$ , we have [21]

$$\nabla_{\zeta_1} \xi = -\zeta_1 + \eta(\zeta_1)\xi, \ g(\xi,\xi) = 1, \text{ and } \eta(\zeta_1) = g(\zeta_1,\xi)$$
 (4)

for any  $\zeta_1$  on  $M^n$ . Next, we have [21]

$$(\nabla_{\zeta_1}\eta)\zeta_2 = -g(\zeta_1,\zeta_2) + \eta(\zeta_1)\eta(\zeta_2), \tag{5}$$

$$K(\zeta_1, \zeta_2)\xi = \eta(\zeta_1)\zeta_2 - \eta(\zeta_2)\zeta_1,$$
(6)

$$K(\zeta_1,\xi)\zeta_2 = g(\zeta_1,\zeta_2)\xi - \eta(\zeta_2)\zeta_1,$$
(7)

$$S(\zeta_1,\xi) = -(n-1)\eta(\zeta_1) \iff Q\xi = -(n-1)\xi,$$
(8)

$$(\pounds_{\xi}g)(\zeta_1,\zeta_2) = 2\{-g(\zeta_1,\zeta_2) + \eta(\zeta_1)\eta(\zeta_2)\},\tag{9}$$

for all  $\zeta_1$ ,  $\zeta_2$  on  $M^n$ . Here, K and Q represent the curvature tensor and the Ricci operator of  $M^n$ , respectively.

**Definition 1.** An  $M^n$  is said to be quasi-Einstein if its  $S(\neq 0)$  satisfies

$$S(\zeta_1,\zeta_2) = \mathfrak{l}g(\zeta_1,\zeta_2) + \mathfrak{m}\eta(\zeta_1)\eta(\zeta_2),$$

where  $\mathfrak{m}$  and  $\mathfrak{l}$  are smooth functions on  $M^n$ . If  $\mathfrak{m} = 0$ , then the manifold is called an Einstein manifold.

**Definition 2.** A partial differential equation  $\Delta u = v$  on a complete  $M^n$  is called a Poisson equation for some smooth functions u and v.

**Remark 1** ([21,22]). An  $(M^3, \tilde{\nabla}, g)$  is a quasi-Einstein manifold of the form

$$S(\zeta_1, \zeta_2) = \left(1 + \frac{R}{2}\right) g(\zeta_1, \zeta_2) - \left(3 + \frac{R}{2}\right) \eta(\zeta_1) \eta(\zeta_2).$$
(10)

**Remark 2** ([21,22]). *In an*  $(M^3, \tilde{\nabla}, g)$ *, we have* 

$$\xi(R) = 2(R+6),$$
 (11)

$$\eta(\nabla_{\xi}\mathcal{D}R) = 4(R+6),\tag{12}$$

where D is the gradient operator of g. From (11), it is noticed that R of  $M^3$  is constant if and only if R = -6.

# 3. $\rho$ -ES on $(M^3, \tilde{\nabla}, g)$

First, we prove the following theorem.

**Theorem 1.** If  $(M^3, \tilde{\nabla}, g)$  admits a  $\rho$ -ES  $(g, E, \sigma, \rho)$ , then its scalar curvature R satisfies the Poisson equation  $\Delta R = \frac{4(2-\sigma-6\rho)}{\rho}$ , provided  $\rho \neq 0$ .

**Proof.** Let the metric of an  $(M^3, \tilde{\nabla}, g)$  be a  $\rho$ -ES  $(g, E, \sigma, \rho)$ , then in view of (10), (2) leads to

$$(\pounds_{E}g)(\zeta_{1},\zeta_{2}) = -2\{1+\sigma+(\frac{1}{2}-\rho)R\}g(\zeta_{1},\zeta_{2}) + (R+6)\eta(\zeta_{1})\eta(\zeta_{2}),$$
(13)

for any vector fields  $\zeta_1$ ,  $\zeta_2$  on  $M^3$ .

Taking covariant derivative of (13) respecting to  $\zeta_3$ , we find

$$(\nabla_{\zeta_3} \pounds_E g)(\zeta_1, \zeta_2) = (\zeta_3 R) \{ (2\rho - 1)g(\zeta_1, \zeta_2) + \eta(\zeta_1)\eta(\zeta_2) \} - (R + 6) \{ g(\zeta_1, \zeta_3)\eta(\zeta_2) + g(\zeta_2, \zeta_3)\eta(\zeta_1) - 2\eta(\zeta_1)\eta(\zeta_2)\eta(\zeta_3) \}.$$
 (14)

As *g* is parallel with respect to  $\nabla$ , then the formula [25]

$$(\pounds_E \nabla_{\zeta_1} g - \nabla_{\zeta_1} \pounds_E g - \nabla_{[E,\zeta_1]} g)(\zeta_2,\zeta_3) = -g((\pounds_E \nabla)(\zeta_1,\zeta_2),\zeta_3) - g((\pounds_E \nabla)(\zeta_1,\zeta_3),\zeta_2)$$

turns to

$$(\nabla_{\zeta_1} \pounds_E g)(\zeta_2, \zeta_3) = g((\pounds_E \nabla)(\zeta_1, \zeta_2), \zeta_3) + g((\pounds_E \nabla)(\zeta_1, \zeta_3), \zeta_2).$$

Since  $\pounds_E \nabla$  is symmetric, therefore we have

$$2g((\pounds_E \nabla)(\zeta_1, \zeta_2), \zeta_3) = (\nabla_{\zeta_1} \pounds_E g)(\zeta_2, \zeta_3) + (\nabla_{\zeta_2} \pounds_E g)(\zeta_1, \zeta_3) - (\nabla_{\zeta_3} \pounds_E g)(\zeta_1, \zeta_2),$$

which in view of (14) gives

$$2g((\pounds_E \nabla)(\zeta_1, \zeta_2), \zeta_3) = (\zeta_1 R) \{ (2\rho - 1)g(\zeta_2, \zeta_3) + \eta(\zeta_2)\eta(\zeta_3) \} \\ + (\zeta_2 R) \{ (2\rho - 1)g(\zeta_1, \zeta_3) + \eta(\zeta_1)\eta(\zeta_3) \} \\ - (\zeta_3 R) \{ (2\rho - 1)g(\zeta_1, \zeta_2) + \eta(\zeta_1)\eta(\zeta_2) \} \\ - 2(R + 6) \{ g(\zeta_1, \zeta_2)\eta(\zeta_3) - \eta(\zeta_1)\eta(\zeta_2)\eta(\zeta_3) \},$$

from which it follows that

$$2(\pounds_{E}\nabla)(\zeta_{1},\zeta_{2}) = (\zeta_{1}R)\{(2\rho-1)\zeta_{2}+\eta(\zeta_{2})\xi\}$$

$$+(\zeta_{2}R)\{(2\rho-1)\zeta_{1}+\eta(\zeta_{1})\xi\}$$

$$-(\mathcal{D}R)\{(2\rho-1)g(\zeta_{1},\zeta_{2})+\eta(\zeta_{1})\eta(\zeta_{2})\}$$

$$-2(R+6)\{g(\zeta_{1},\zeta_{2})\xi-\eta(\zeta_{1})\eta(\zeta_{2})\xi\}.$$
(15)

Replacing  $\zeta_2$  by  $\xi$  and  $\zeta_1$  by  $\zeta_2$  in (15), we have

$$(\pounds_E \nabla)(\zeta_2, \xi) = \rho_g(\mathcal{D}R, \zeta_2)\xi - \rho(\mathcal{D}R)\eta(\zeta_2) + (R+6)\{(2\rho-1)\zeta_2 + \eta(\zeta_2)\xi\}.$$
 (16)

The covariant differentiation of (16) respecting to  $\zeta_1$  yields

$$\begin{aligned} (\nabla_{\zeta_{1}} \pounds_{E} \nabla)(\zeta_{2}, \xi) &= 2(\zeta_{1} R) \{ (2\rho - 1)\zeta_{2} + \eta(\zeta_{2})\xi \} \\ &+ (\zeta_{2} R) \{ (\rho - 1)\zeta_{1} + \eta(\zeta_{1})\xi \} \\ &- (\mathcal{D} R) \{ (\rho - 1)g(\zeta_{1}, \zeta_{2}) + \eta(\zeta_{1})\eta(\zeta_{2}) \} \\ &- 3(R + 6) \{ g(\zeta_{1}, \zeta_{2})\xi - \eta(\zeta_{1})\eta(\zeta_{2})\xi \} \\ &- (R + 6) \{ (2\rho - 1)\eta(\zeta_{1})\zeta_{2} + \eta(\zeta_{2})\zeta_{1} \} \\ &+ \rho_{g} (\nabla_{\zeta_{1}} \mathcal{D} R, \zeta_{2})\xi - \rho(\nabla_{\zeta_{1}} \mathcal{D} R)\eta(\zeta_{2}), \end{aligned}$$
(17)

where (4), (5) and (16) being used.

Again from [25], we have

$$(\pounds_E K)(\zeta_1, \zeta_2)\zeta_3 = (\nabla_{\zeta_1} \pounds_E \nabla)(\zeta_2, \zeta_3) - (\nabla_{\zeta_2} \pounds_E \nabla)(\zeta_1, \zeta_3), \tag{18}$$

which by putting  $\zeta_3 = \xi$  and using (17) becomes

$$(\pounds_{E}K)(\zeta_{1},\zeta_{2})\xi = g(\mathcal{D}R,\zeta_{1})\{(3\rho-1)\zeta_{2}+\eta(\zeta_{2})\xi\}$$

$$-g(\mathcal{D}R,\zeta_{2})\{(3\rho-1)\zeta_{1}+\eta(\zeta_{1})\xi\}$$

$$+2(R+6)(\rho-1)\{\eta(\zeta_{2})\zeta_{1}-\eta(\zeta_{1})\zeta_{2}\}$$

$$+\rho g(\nabla_{\zeta_{1}}\mathcal{D}R,\zeta_{2})\xi -\rho g(\nabla_{\zeta_{2}}\mathcal{D}R,\zeta_{1})\xi$$

$$-\rho(\nabla_{\zeta_{1}}\mathcal{D}R)\eta(\zeta_{2}) +\rho(\nabla_{\zeta_{2}}\mathcal{D}R)\eta(\zeta_{1}).$$

$$(19)$$

Contracting (19) respecting to  $\zeta_1$  then using (4) and (11) we lead to

$$(\pounds_E S)(\zeta_2,\xi) = (1-6\rho)\zeta_2(R) + 2(R+6)(2\rho-1)\eta(\zeta_2) + \rho g(\nabla_{\xi} \mathcal{D}R,\zeta_2)\xi - \rho(\Delta R)\eta(\zeta_2).$$
 (20)

By putting  $\zeta_2 = \xi$  in (20) then using (4), (11) and (12), we find

$$(\pounds_E S)(\xi,\xi) = -4\rho(R+6) - \rho(\Delta R).$$
(21)

The Lie derivative of (8) respecting to *E* leads to

$$(\pounds_E S)(\xi,\xi) = 4\eta(\pounds_E \xi). \tag{22}$$

Putting  $\zeta_1 = \zeta_2 = \xi$  in (13) infers

$$(\pounds_E g)(\xi,\xi) = -2\sigma + 2\rho R + 4. \tag{23}$$

The Lie derivative of  $g(\xi, \xi) = 1$  gives

$$(\pounds_E g)(\xi,\xi) = -2\eta(\pounds_E \xi).$$
(24)

Now combining (21)–(25) we deduce

$$\Delta R = \frac{4(2 - \sigma - 6\rho)}{\rho}, \text{ provided } \rho \neq 0.$$
(25)

This completes the proof.  $\Box$ 

It is well-known that the  $\rho$ -ES Equation (2) on  $M^n$  with the soliton constant  $\rho = \frac{1}{2}, \frac{1}{n}, \frac{1}{2(n-1)}$  reduces to the Einstein soliton, traceless Ricci soliton, Schouten soliton, respectively. It is also known that a smooth function  $\mathfrak{f}$  on an  $M^n$  is called harmonic, sub-harmonic or superharmonic if  $\Delta \mathfrak{f} = 0$ ,  $\geq 0$  or  $\leq 0$ , respectively. These facts together with Theorem 1 state the following:

**Corollary 1.** Let  $(M^3, \tilde{\nabla}, g)$  admit a  $\rho$ -ES, then we have

| Value of p           | Solitons                | Poisson equation                      | Condition for R to be subharmonic and superharmonic   |
|----------------------|-------------------------|---------------------------------------|---|
| $\rho = \frac{1}{2}$ | Einstein soliton        | $\Delta R = -8(\sigma + 1)$           | ( <i>i</i> ) <i>R</i> is subharmonic if $\sigma \leq -1$ ,<br>( <i>ii</i> ) <i>R</i> is superharmonic if $\sigma \geq -1$ ,                   |
| $\rho = \frac{1}{3}$ | traceless Ricci soliton | $\Delta R = -12\sigma$                | ( <i>i</i> ) <i>R</i> is subharmonic if $\sigma \leq 0$ ,<br>( <i>ii</i> ) <i>R</i> is superharmonic if $\sigma \geq 0$ ,                     |
| $\rho = \frac{1}{4}$ | Schouten soliton        | $\Delta R = 16(\frac{1}{2} - \sigma)$ | ( <i>i</i> ) <i>R</i> is subharmonic if $\sigma \leq \frac{1}{2}$ ,<br>( <i>ii</i> ) <i>R</i> is superharmonic if $\sigma \geq \frac{1}{2}$ . |

**Remark 3.** The  $\rho$ -ES on an  $M^n$  with  $\rho = 0$  reduces to the RS. The properties of RS on  $(M^3, \tilde{\nabla}, g)$  have been explored by Chaubey and De [22]. Thus, we can say that the Theorem 1 generalizes the study of Einstein soliton, traceless RS and the Schouten soliton on  $(M^3, \tilde{\nabla}, g)$ .

It is well-known that the Poisson equation  $\Delta u = v$  with v = 0 becomes a Laplace equation. Suppose that an  $(M^3, \tilde{\nabla}, g)$  does not admit RS. Then, Theorem 1 and above discussion state:

**Corollary 2.** If  $(M^3, \tilde{\nabla}, g)$  admits a  $\rho$ -ES, which is not a RS  $(\rho \neq 0)$ , then R of  $M^3$  satisfies Laplace equation if and only if  $\sigma = 2(1 - 3\rho)$ .

Let  $(M^3, \tilde{\nabla}, g)$  admit a  $\rho$ -ES. If *R* of  $M^3$  satisfies the Laplace equation, then  $\sigma = 2(1 - 3\rho)$ . The  $\rho$ -ES under consideration to be steady, shrinking or expanding if  $\rho$  is equal to, less than or greater than  $\frac{1}{3}$ . Thus, we write our corollary as

**Corollary 3.** Let the metric of an  $(M^3, \tilde{\nabla}, g)$  be  $\rho$ -ES, which is not a RS  $(\rho \neq 0)$ . If R of  $M^3$  satisfies the Laplace equation, then the  $\rho$ -ES is steady, shrinking or expanding if  $\rho = \frac{1}{3}$ ,  $\rho < \frac{1}{3}$  or  $\rho > \frac{1}{3}$ , respectively.

# 4. Gradient $\rho$ -ES on $(M^3, \tilde{\nabla}, g)$

**Theorem 2.** Let  $(M^3, \tilde{\nabla}, g)$  admit a gradient  $\rho$ -ES. Then, either  $M^3$  is Einstein or the gradient  $\rho$ -ES is steady type gradient traceless RS.

**Proof.** Let the metric of an  $(M^3, \tilde{\nabla}, g)$  be a gradient  $\rho$ -ES. Then, (3) can be written as

$$\nabla_{\zeta_1} \mathcal{DF} + Q\zeta_1 + (\sigma - \rho R)\zeta_1 = 0, \tag{26}$$

for all  $\zeta_1$  on  $M^3$ .

The covariant differentiation of (26) with respect to  $\zeta_2$  leads to

$$\nabla_{\zeta_2} \nabla_{\zeta_1} \mathcal{DF} = -(\nabla_{\zeta_2} Q)\zeta_1 - Q(\nabla_{\zeta_2} \zeta_1) - (\sigma - \rho R)\nabla_{\zeta_2} \zeta_1 + \rho \zeta_2(R)\zeta_1.$$
(27)

Interchanging  $\zeta_1$  and  $\zeta_2$  in (27) leads to

$$\nabla_{\zeta_1} \nabla_{\zeta_2} \mathcal{DF} = -(\nabla_{\zeta_1} Q)\zeta_2 - Q(\nabla_{\zeta_1} \zeta_2) - (\sigma - \rho R)\nabla_{\zeta_1} \zeta_2 + \rho\zeta_1(R)\zeta_2.$$
(28)

By plugging of (26)–(28), we find

$$K(\zeta_1,\zeta_2)\mathcal{DF} = -(\nabla_{\zeta_1}Q)\zeta_2 + (\nabla_{\zeta_2}Q)\zeta_1 + \rho\{\zeta_1(R)\zeta_2 - \zeta_2(R)\zeta_1\}.$$

Contracting the forgoing equation along  $\zeta_1$ , we obtain

$$S(\zeta_2, \mathcal{DF}) = \frac{(1-4\rho)}{2}\zeta_2(R).$$
<sup>(29)</sup>

In account of (10), we have

$$S(\zeta_2, \mathcal{DF}) = (1 + \frac{R}{2})\zeta_2(\mathcal{F}) - (3 + \frac{R}{2})\eta(\zeta_2)\xi(\mathcal{F}).$$
(30)

Thus, from (29) and (30), it follows that

$$(1-4\rho)\zeta_2(R) = (R+2)\zeta_2(\mathcal{F}) - (R+6)\eta(\zeta_2)\xi(\mathcal{F}).$$
(31)

By putting  $\zeta_2 = \xi$  in (31), then using (4) and (11), we find

$$\xi(\mathcal{F}) = -\frac{1}{2}(1 - 4\rho)(R + 6).$$
(32)

By using (32) and (31) turns to

$$(1-4\rho)\zeta_2(R) = (R+2)\zeta_2(\mathcal{F}) + \frac{1}{2}(R+6)^2(1-4\rho)\eta(\zeta_2).$$
(33)

The covariant differentiation of (33) along  $\zeta_1$  leads to

$$(1-4\rho)g(\nabla_{\zeta_{1}}\mathcal{D}R,\zeta_{2}) = \zeta_{1}(R)\zeta_{2}(\mathcal{F}) + (R+2)g(\nabla_{\zeta_{1}}\mathcal{D}\mathcal{F},\zeta_{2}) + (R+6)(1-4\rho)\zeta_{1}(R)\eta(\zeta_{2}) + \frac{1}{2}(R+6)^{2}(1-4\rho)\{\eta(\zeta_{1})\eta(\zeta_{2}) - g(\zeta_{1},\zeta_{2})\}.$$
(34)

Interchanging  $\zeta_1$  and  $\zeta_2$  in (34), we have

$$(1-4\rho)g(\nabla_{\zeta_{2}}\mathcal{D}R,\zeta_{1}) = \zeta_{2}(R)\zeta_{1}(\mathcal{F}) + (R+2)g(\nabla_{\zeta_{2}}\mathcal{D}\mathcal{F},\zeta_{1})$$

$$+(R+6)(1-4\rho)\zeta_{2}(R)\eta(\zeta_{1})$$

$$+\frac{1}{2}(R+6)^{2}(1-4\rho)\{\eta(\zeta_{1})\eta(\zeta_{2})-g(\zeta_{1},\zeta_{2})\}.$$
(35)

Equating the left hand sides of last two equations gives

$$\zeta_1(R)\zeta_2(\mathcal{F}) + (R+6)(1-4\rho)\zeta_1(R)\eta(\zeta_2) -\zeta_2(R)\zeta_1(\mathcal{F}) - (R+6)(1-4\rho)\zeta_2(R)\eta(\zeta_1) = 0,$$

which by replacing  $\zeta_2 = \xi$  then using (4), (11) and (32) takes the form

$$(R+6)\{(1-4\rho)\zeta_1(R)-4\zeta_1(\mathcal{F})-4(R+6)(1-4\rho)\eta(\zeta_1)\}=0.$$

Thus, we have either R = -6, or  $(1 - 4\rho)\zeta_1(R) = 4\zeta_1(\mathcal{F}) + 4(R + 6)(1 - 4\rho)\eta(\zeta_1)$ . If we firstly suppose that  $R \neq -6$  and  $(1 - 4\rho)\zeta_1(R) = 4\zeta_1(\mathcal{F}) + 4(R + 6)(1 - 4\rho)\eta(\zeta_1)$ , which by virtue of (33) turns to

$$(R-2)\{2\zeta_1(\mathcal{F}) + (R+6)(1-4\rho)\eta(\zeta_1)\} = 0,$$
(36)

which refers that either R = 2 or  $\zeta_1(\mathcal{F}) = -\frac{1}{2}(R+6)(1-4\rho)\eta(\zeta_1)$ . From (11), it is obvious that if R is constant, then its value must be -6, which shows that R = 2 is inadmissible. Thus, we have  $\zeta_1(\mathcal{F}) = -\frac{1}{2}(R+6)(1-4\rho)\eta(\zeta_1)$ , which is equivalent to

$$\mathcal{DF} = -\frac{1}{2}(R+6)(1-4\rho)\xi = \xi(\mathcal{F})\xi.$$
(37)

Thus, the gradient of  $\mathcal{F}$  is pointwise collinear with  $\xi$ . Now, taking the covariant derivative of (37) with respect to  $\zeta_1$  and using (4), we have

$$\nabla_{\zeta_1} \mathcal{DF} = \zeta_1(\xi(\mathcal{F}))\xi - \xi(\mathcal{F})(\zeta_1 - \eta(\zeta_1)\xi).$$
(38)

Therefore, from (26) and (38), we obtain

$$Q\zeta_1 + (\sigma - \rho R)\zeta_1 = -\zeta_1(\xi(\mathcal{F}))\xi + \xi(\mathcal{F})(\zeta_1 - \eta(\zeta_1)\xi).$$
(39)

Now, by replacing  $\zeta_1$  by  $\xi$  in (39) then using (8), (11) and (32) we lead to

$$\sigma = (1 - 3\rho)(R + 8). \tag{40}$$

Let us suppose that  $\rho = \frac{1}{3}$ , that is, the gradient  $\rho$ -ES on an  $M^3$  is gradient traceless RS. This fact together with Equation (40) leads to  $\sigma = 0$ . Thus, the gradient traceless RS is steady. This completes the proof.  $\Box$ 

**Theorem 3.** Let an  $(M^3, \tilde{\nabla}, g)$  be a non-gradient traceless RS. Then, the gradient  $\rho$ -ES is trivial soliton with constant  $\sigma = 2(1 - 3\rho)$ . Also, the  $\rho$ -ES is shrinking and expanding according to  $\rho > \frac{1}{3}$  and  $\rho < \frac{1}{3}$ .

**Proof.** Now, we suppose that  $\rho \neq \frac{1}{3}$ . Thus, (40) leads to

$$R = \frac{\sigma}{1 - 3\rho} - 8,\tag{41}$$

which informs that *R* is constant and hence (11) infers that R = -6. This contradicts our hypothesis  $R \neq -6$ .

Secondly, we consider that R = -6 and  $(1 - 4\rho)\zeta_1(R) \neq 4\zeta_1(\mathcal{F}) + 4(R + 6)(1 - 4\rho)\eta(\zeta_1)$ . For R = -6, (33) informs that  $\mathcal{F} \in \mathcal{R}$  and hence the GRBS on the manifold is trivial. Moreover, the Riemannian 3-manifold under assumption is an Einstein manifold with  $\sigma = 2(1 - 3\rho)$ . This completes the proof.  $\Box$ 

Let us suppose that an  $(M^3, \tilde{\nabla}, g)$  admits a proper gradient  $\rho$ -ES. Then, the  $\rho$ -ES reduces to the gradient traceless RS and  $\rho = \frac{1}{3}, \sigma = 0$ . Using these facts in (26) and then contracting the foregoing equation over  $\zeta_1$  gives  $\Delta \mathcal{F} = 0$ .

A smooth function  $\mathfrak{h}$  on an  $M^n$  is called harmonic if  $\Delta \mathfrak{h} = 0$ .

The above discussions state the following:

**Corollary 4.** Let a complete  $(M^3, \tilde{\nabla}, g)$  admit a proper gradient  $\rho$ -ES. Then the gradient function of the gradient  $\rho$ -ES is harmonic.

Contracting (38) over  $\zeta_1$ , we find

$$\Delta \mathcal{F} = \xi(\xi(\mathcal{F})) - 2\xi(\mathcal{F}).$$

Again, considering  $\sigma = 0$ ,  $\rho = \frac{1}{3}$  and then contracting (26) over  $\zeta_1$ , we conclude that

$$\Delta \mathcal{F} = 0.$$

The last two equations show that  $\xi(\xi(\mathcal{F})) - 2\xi(\mathcal{F}) = 0$ . Let  $\xi = \frac{\partial}{\partial t}$ . Thus, we notice that the potential function  $\mathcal{F}$  satisfies the PDE

$$\frac{\partial^2 \mathcal{F}}{\partial t^2} - 2\frac{\partial \mathcal{F}}{\partial t} = 0$$

It is obvious that  $\mathcal{F} = Ae^{2t} + B$  for smooth functions *A* and *B*, which are independent of *t*, is the solution of the above PDE. Now, we list our results in the following:

**Corollary 5.** Let the metric of a complete  $(M^3, \tilde{\nabla}, g)$  admit a proper gradient  $\rho$ -ES. Then, the potential function  $\mathcal{F}$  of such soliton satisfies the PDE  $\frac{\partial^2 \mathcal{F}}{\partial t^2} - 2\frac{\partial \mathcal{F}}{\partial t} = 0$ , and it can be evaluated by  $\mathcal{F} = Ae^{2t} + B$ .

#### 5. Example

We consider the manifold  $M^3 = \{(w_1, w_2, w_3) \in \mathbb{R}^3\}$ , where  $(w_1, w_2, w_3)$  are the usual coordinates in  $\mathbb{R}^3$ . Let  $u_1, u_2$  and  $u_3$  be the vector fields on  $M^3$  given by

$$u_1 = e^{\mathfrak{b}w_3 + w_1} \frac{\partial}{\partial w_1}, \quad u_2 = e^{\mathfrak{b}w_3 + w_2} \frac{\partial}{\partial w_2}, \quad u_3 = \frac{1}{\mathfrak{b}} \frac{\partial}{\partial w_3} = \xi,$$

where  $\mathfrak{b}(\neq 0) \in \mathcal{R}$ . Then,  $\{u_1, u_2, u_3\}$  forms a basis in the module of the vector fields of  $M^3$ .

Let the Riemannian metric *g* be defined by

$$g(u_p, u_q) = \begin{cases} 1, & 1 \le p = q \le 3, \\ 0, & otherwise. \end{cases}$$

Hence,  $M^3$  is a Riemannian manifold of dimension 3. Let the 1-form  $\eta$  on  $M^3$  be defined by  $\eta(\zeta_1) = g(\zeta_1, u_3) = g(\zeta_1, \zeta)$  for all  $\zeta_1$  on  $M^3$ . Now, by direct computations, we obtain

$$[u_1, u_2] = 0, \ [u_1, u_3] = -u_1, \ [u_2, u_3] = -u_2.$$

By using Koszul's formula, we obtain

$$\nabla_{u_p} u_q = \begin{cases} -u_p, & p = 1, 2, q = 3, \\ u_3, & 1 \le p = q \le 2, \\ 0, & otherwise. \end{cases}$$

Now we suppose that  $\zeta_1 = \zeta_1^1 u_1 + \zeta_1^2 u_2 + \zeta_1^3 u_3$ , then for  $\xi = u_3$  it follows that  $\nabla_{\zeta_1} \xi = -\zeta_1 + \eta(\zeta_1)\xi$ . It can be easily seen that  $\tilde{\nabla}$  defined on  $M^3$  satisfies the conditions

$$ilde{T}(\zeta_1,\zeta_2) = -\eta(\zeta_1)\zeta_2 + \eta(\zeta_2)\zeta_1, \ ilde{
abla}g = 0, \ ext{and} \ ilde{
abla}\xi = 0,$$

for arbitrary vector fields  $\zeta_1$  and  $\zeta_2$  on  $M^3$ , where  $\tilde{T}$  indicates the torsion tensor of  $\tilde{\nabla}$ . Thus, we can say that  $\tilde{\nabla}$  is a semi-symmetric metric  $\xi$ -connection on  $M^3$ .

The non-zero constituents of *K* are obtained as follows:

$$K(u_1, u_3)u_1 = u_3, \quad K(u_1, u_2)u_1 = u_2, \quad K(u_2, u_3)u_2 = u_3,$$
  
 $K(u_1, u_2)u_2 = K(u_1, u_3)u_3 = -u_1, \quad K(u_2, u_3)u_3 = -u_2.$ 

By using above components of the curvature tensor *K* we obtain

$$S(u_p, u_q) = -2, \ 1 \le p = q \le 3,$$

from which we obtain R = -6.

Now, by taking  $\mathcal{DF} = (u_1\mathcal{F})u_1 + (u_2\mathcal{F})u_2 + (u_3\mathcal{F})u_3$ , we have

$$\begin{aligned} \nabla_{u_1} \mathcal{DF} &= (u_1(u_1\mathcal{F}) - u_3\mathcal{F})u_1 + (u_1(u_2\mathcal{F}))u_2 + (u_1(u_3\mathcal{F}) + u_1\mathcal{F})u_3, \\ \nabla_{\mathcal{E}_2} \mathcal{DF} &= (u_2(u_1\mathcal{F}))u_1 + (u_2(u_2\mathcal{F}) - u_3\mathcal{F})u_2 + (u_2(u_3\mathcal{F}) + \mathcal{F}_2\mathcal{F})\mathcal{F}_3, \\ \nabla_{\mathcal{E}_3} \mathcal{DF} &= (u_3(u_1\mathcal{F}))u_1 + (u_3(u_2\mathcal{F}))u_2 + (u_3(u_3\mathcal{F}))u_3. \end{aligned}$$

Thus, by virtue of (26), we obtain

$$\begin{cases} u_{1}(u_{1}\mathcal{F}) - u_{3}\mathcal{F} = 2 - 6\rho - \sigma, \\ u_{2}(u_{2}\mathcal{F}) - u_{3}\mathcal{F} = 2 - 6\rho - \sigma, \\ u_{3}(u_{3}\mathcal{F}) = 2 - 6\rho - \sigma, \\ u_{1}(u_{2}\mathcal{F}) = 0, \\ u_{2}(u_{1}\mathcal{F}) = 0, \\ u_{2}(u_{3}\mathcal{F}) + u_{2}\mathcal{F} = 0. \end{cases}$$

$$(42)$$

Thus, the relations in (42) are, respectively, amounting to

$$\begin{split} e^{2(\mathfrak{b}w_3+w_1)} \Big[ \frac{\partial^2 \mathcal{F}}{\partial w_1^2} + \frac{\partial \mathcal{F}}{\partial w_1} \Big] &- \frac{1}{\mathfrak{b}} \frac{\partial \mathcal{F}}{\partial w_3} = 2 - 6\rho - \sigma, \\ e^{2(\mathfrak{b}w_3+w_1)} \Big[ \frac{\partial^2 \mathcal{F}}{\partial w_2^2} + \frac{\partial \mathcal{F}}{\partial w_2} \Big] &- \frac{1}{\mathfrak{b}} \frac{\partial \mathcal{F}}{\partial w_3} = 2 - 6\rho - \sigma, \\ \frac{1}{\mathfrak{b}^2} \frac{\partial^2 \mathcal{F}}{\partial w_3^2} = 2 - 6\rho - \sigma, \\ \frac{\partial^2 \mathcal{F}}{\partial w_1 \partial w_2} = 0, \\ \frac{\partial^2 \mathcal{F}}{\partial w_2 \partial w_1} = 0, \\ \frac{1}{\mathfrak{b}} \Big[ \frac{\partial^2 \mathcal{F}}{\partial w_2 \partial w_3} + \frac{\partial \mathcal{F}}{\partial w_2} \Big] = 0. \end{split}$$

From the above relations, it is noticed that  $\mathcal{F} \in \mathcal{R}$  for  $\sigma = 2 - 6\rho$ . Hence, the Equation (26) is satisfied. Thus, *g* is a gradient  $\rho$ -ES with the soliton vector field  $E = \mathcal{DF}$ , where  $\mathcal{F} \in \mathcal{R}$  and  $\sigma = 2 - 6\rho$ . For  $\rho = \frac{1}{3}$ , we obtain  $\sigma = 0$ , i.e., the gradient  $\rho$ -ES is trivial with constant  $\sigma = 2 - 6\rho$ . Thus, Theorem 2 is verified.

## 6. Results and Discussion

It is well known that the  $\rho$ -Einstein soliton Equation (2) with  $\rho = 0$  becomes the Ricci soliton equation, which has been studied in [22]. Thus, we can say that the  $\rho$ -Einstein soliton is a natural generalization of Ricci soliton. In this manuscript, we have explored the properties of  $\rho$ -Einstein solitons in Riemannian geometry, which generalizes the results of [22].

# 7. Conclusions

To prove the curvatures invariant, Chauey et al. [23] defined the notion of semisymmetric metric *P*-connection in Riemannian setting, which is a particular case of Riemannian concircular structure manifold [26]. This topic has great applications in differential equations. We proved that the scalar curvature of Riemannian 3-manifolds endowed with a semi-symmetric metric  $\zeta$ -connection and Ricci–Bourguignon soliton satisfies the Poisson and Laplace equations. It is well known that the Poisson and Laplace equations play a crucial role in the development of engineering, physics, mathematics, etc. We have also established the conditions for which the scalar curvature is harmonic, sub-harmonic and super-harmonic. We also established the existence condition of a gradient  $\rho$ -Einstein soliton in the Riemannian 3-manifolds, and consequently we proved some results. To verify our results, we constructed a non-trivial example of a three-dimensional Riemannian manifold equipped with a semi-symmetric metric  $\zeta$ -connection. These topics are modern and have a lot of scope for researchers.

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#### References

- 1. Hamilton, R.S. The Ricci Flow on Surfaces, Mathematics and General Relativity. *Contemp. Math.* **1988**, *71*, 237–262.
- 2. Besse, A.L. Einstein Manifolds, Classics in Mathematics; Springer: Berlin/Heidelberg, Germany, 2008.
- 3. Bourguignon, J.P. Ricci curvature and Einstein metrics. Glob. Differ. Geom. Glob. Anal. Lect. Notes Math. 1981, 838, 42–63.
- Bourguignon, J.P.; Lawson, H.B. Stability and isolation phenomena for Yang-mills fields. *Commun. Math. Phys.* 1981, 79, 189–230. [CrossRef]
- 5. Dwivedi, S. Some results on Ricci-Bourguignon solitons and almost solitons. Can. Math. Bull. 2020, 64, 1–15. [CrossRef]
- 6. Huang, G. Integral pinched gradient shrinking *ρ*-Einstein solitons. J. Math. Anal. Appl. 2017, 451, 1045–1055. [CrossRef]
- Mondal, C.K.; Shaikh, A.A. Some results on η-Ricci Soliton and gradient ρ-Einstein soliton in a complete Riemannian manifold. *Commun. Korean Math. Soc.* 2019, 34, 1279–1287.
- 8. Patra, D.S. Some characterizations of *ρ*-Einstein solitons on Sasakian manifolds. *Can. Math. Bull.* **2022**, *65*, 1036–1049. [CrossRef]
- 9. Shaikh, A.A.; Cunha, A.W.; Mandal, P. Some characterizations of  $\rho$ -Einstein solitons. J. Geom. Phys. **2021**, 166, 104270. [CrossRef]
- Shaikh, A.A.; Mandal, P.; Mondal, C.K. Diameter estimation of gradient *ρ*-Einstein solitons. *J. Geom. Phys.* 2022, 177, 104518.
   [CrossRef]
- 11. Suh, Y.J. Ricci-Bourguignon solitons on real hypersurfaces in the complex hyperbolic quadric. *Rev. Real Acad. Cienc. Exactas Fis. Nat. Ser. A-Mat.* **2022**, *116*, 110. [CrossRef]
- Venkatesha, V.; Kumara, H.A. Gradient *ρ*-Einstein soliton on almost Kenmotsu manifolds. *Ann. Dell' Univ. Ferrara* 2019, 65, 375–388. [CrossRef]
- 13. Blaga, A.M. Some geometrical aspects of Einstein, Ricci and Yamabe solitons. J. Geom. Symmetry Phys. 2019, 52, 17–26. [CrossRef]
- 14. Chen, Z.; Li, Y.; Sarkar, S.; Dey, S.; Bhattacharyya, A. Ricci soliton and certain related metrics on a three-dimensional trans-Sasakian manifold. *Universe* **2022**, *8*, 595. [CrossRef]

- Li, Y.; Haseeb, A.; Ali, M. LP-Kenmotsu manifolds admitting η-Ricci-Yamabe Solitons and spacetime. J. Math. 2022, 2022, 6605127.
   [CrossRef]
- 16. Suh, Y.J. Yamabe and gradient Yamabe solitons in the complex hyperbolic two-plane Grassmannians. *Rev. Math. Phys.* **2022**, 34, 2250024. [CrossRef]
- 17. Suh, Y.J. Yamabe and quasi-Yamabe solitons on hypersurfaces in the complex hyperbolic space. *Mediterr. J. Math.* **2023**, *20*, 69. [CrossRef]
- 18. Turki, N.B.; Blaga, A.M.; Deshmukh, S. Soliton-type equations on a Riemannian manifold. *Mathematics* 2022, 10, 633. [CrossRef]
- Yoldas, H.I. On Kenmotsu manifolds admitting η-Ricci-Yamabe solitons. *Int. J. Geom. Methods Mod. Phys.* 2021, 18, 2150189. [CrossRef]
- 20. Yano, K. On semi-symmetric metric connections. Rev. Roumaine Math. Pures Appl. 1970, 15, 1579–1586.
- 21. Chaubey, S.K.; De, U.C. Characterization of three-dimensional Riemannian manifolds with a type of semi-symmetric metric connection admitting Yamabe soliton. *J. Geom. Phys.* **2020**, *157*, 103846. [CrossRef]
- Chaubey, S.K.; De, U.C. Three dimensional Riemannian manifolds and Ricci solitons. *Quaest. Math.* 2022, 45, 765–778. [CrossRef]
   Chaubey, S.K.; Lee, J.W.; Yadav, S. Riemannian manifolds with a semisymmetric metric *P*-connection. *J. Korean Math. Soc.* 2019,
- 56, 1113–1129.
- Haseeb, A.; Chaubey, S.K.; Khan, M.A. Riemannian 3-manifolds and Ricci-Yamabe solitons. Int. J. Geom. Methods Mod. Phys. 2023, 20, 2350015. [CrossRef]
- 25. Yano, K. Integral Formulas in Riemannian Geometry, Pure and Applied Mathematics; Marcel Dekker: New York, NY, USA, 1970; Volume I.
- 26. Chaubey, S.K.; Suh, Y.J. Riemannian concircular structure manifolds. Filomat 2022, 36, 6699–6711. [CrossRef]

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