# A Solitonic Study of Riemannian Manifolds Equipped with a Semi-Symmetric Metric $\boldsymbol{\xi}$-Connection 

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#### Abstract

The aim of this paper is to characterize a Riemannian 3-manifold $M^{3}$ equipped with a semisymmetric metric $\xi$-connection $\tilde{\nabla}$ with $\rho$-Einstein and gradient $\rho$-Einstein solitons. The existence of a gradient $\rho$-Einstein soliton in an $M^{3}$ admitting $\tilde{\nabla}$ is ensured by constructing a non-trivial example, and hence some of our results are verified. By using standard tensorial technique, we prove that the scalar curvature of $\left(M^{3}, \tilde{\nabla}\right)$ satisfies the Poisson equation $\Delta R=\frac{4(2-\sigma-6 \rho)}{\rho}$.


Keywords: Riemannian manifolds; $\rho$-Einstein solitons; Einstein manifolds; Poisson equation
MSC: 53E20; 53C25; 53C21

## 1. Introduction

The Ricci and other geometric flows are active topics of current research in mathematics, physics and engineering. The Ricci flow [1] is defined on a Riemannian $n$-manifold $\left(M^{n}, g\right)$ by an evolution equation for metric $g(t)$ of the form $\frac{\partial g}{\partial t}=-2 S$, where $S$ is the Ricci tensor of $M^{n}$ and $t$ indicates the time. The metric $g$ on $M^{n}$ satisfies the Ricci soliton (in short, RS) equation $£_{E} g+2 \sigma g+2 S=0$, where $E$ is a vector field on $M^{n}, \sigma \in \mathcal{R}$ (the set of real numbers), and $£_{E}$ represents the Lie derivative operator in the direction of $E$ on $M^{n}$. A RS is called expanding (steady or shrinking) if $\sigma>0(\sigma=0$ or $\sigma<0)$. If $E=0$ or Killing, then the RS is called a trivial RS, and $M^{n}$ becomes an Einstein manifold. Thus the RS is a basic generalization of an Einstein manifold [2]. If $\mathcal{F}$ is a smooth function such that $E=\mathcal{D} \mathcal{F}$ for the gradient operator $\mathcal{D}$ of $g$, then the RS is described as a gradient Ricci soliton (GRS), $E$ is referred to as the potential vector field, and $\mathcal{F}$ is called the potential function. Thus, the RS equation becomes Hess $\mathcal{F}+\sigma g+S=0$, where Hess $\mathcal{F}$ is the Hessian of $\mathcal{F}$ and $($ Hess $\mathcal{F})\left(\zeta_{1}, \zeta_{2}\right)=g\left(\nabla_{\zeta_{1}} D \mathcal{F}, \zeta_{2}\right)$ for all vector fields $\zeta_{1}$ and $\zeta_{2}$ on $M^{n}$. Here, $\nabla$ stands for the Levi-Civita connection.

The notion of Ricci-Bourguignon flow, a natural generalization of Ricci flow, has been proposed in [3] and is described on an $M^{n}$ as:

$$
\begin{equation*}
\frac{\partial g}{\partial t}=-2(S-\rho R g), \quad g(0)=g_{0} \tag{1}
\end{equation*}
$$

where $R$ is the scalar curvature and $\rho \in \mathcal{R}$. It is to be noticed that for the specific values of $\rho$, the following cases for the tensor $S-\rho R g$ appeared in (1) [4] are obtained:
(i) $\rho=\frac{1}{2}$, the Einstein tensor $S-\frac{R}{2} g$, (for Einstein soliton),
(ii) $\rho=\frac{1}{n}$, the trace-less Ricci tensor $S-\frac{R}{n} g$,
(iii) $\rho=\frac{1}{2(n-1)}$, the Schouten tensor $S-\frac{R}{2(n-1)} g$, (for Schouten soliton),
(iv) $\rho=0$, the Ricci tensor $S$ (for RS).

An $\left(M^{n}, g\right), n \geq 3$ is said to be a $\rho$-Einstein soliton (or $\rho$-ES) $(g, E, \sigma, \rho)$ if

$$
\begin{equation*}
£_{E} g+2 S+2(\sigma-\rho R) g=0 . \tag{2}
\end{equation*}
$$

Similar to the RS, a $\rho$-ES is called expanding (steady or shrinking) if $\sigma>0(\sigma=0$ or $\sigma<0)$. If $E=\mathcal{D} \mathcal{F}$, then $\left(M^{n}, g\right)$ is called a gradient $\rho$-Einstein soliton (or gradient $\rho$-ES). Hence, (2) takes the form

$$
\begin{equation*}
H e s s \mathcal{F}+S+(\sigma-\rho R) g=0 \tag{3}
\end{equation*}
$$

where Hess $\mathcal{F}$ denotes the Hessian of $\mathcal{F} \in C^{\infty}\left(M^{n}\right)$ and defined by Hess $\mathcal{F}=\nabla \nabla \mathcal{F}$. Recently, $\rho$-Einstein solitons have been studied by several authors, such as [5-12]. On the other hand, we recommend the papers [13-19] for the studies of Ricci, Yamabe, RicciYamabe, $\eta$-Ricci-Yamabe and quasi-Yamabe solitons on different geometric structures.

In this paper, we have made an effort to the solitonic study of a 3-dimensional Riemannian manifold $M^{3}$ equipped with a semi-symmetric metric $\xi$-connection $\tilde{\nabla}$. To achieve the goal, we present our work as follows: In Section 2, we gather the basic information of a Riemannian 3-manifold equipped with a semi-symmetric metric $\xi$-connection $\left(M^{3}, \tilde{\nabla}, g\right)$, definitions and Lemmas. The properties of $\rho$-ES in $\left(M^{3}, \tilde{\nabla}, g\right)$ are studied in Section 3. We address the properties of gradient $\rho$-ES in $\left(M^{3}, \tilde{\nabla}, g\right)$ in Section 4. In the last section, we model a non-trivial example of $\left(M^{3}, \tilde{\nabla}, g\right)$ admitting a gradient $\rho$-ES, and prove our results.

## 2. Riemannian Manifolds with a Semi-Symmetric Metric $\boldsymbol{\xi}$-Connection

In 1970, Yano [20] investigated the properties of a semi-symmetric metric connection $\tilde{\nabla}$ on Riemannian $n$-manifolds $M^{n}$ and defined by $\tilde{\nabla}_{\zeta_{1} \zeta_{2}}=\nabla_{\zeta_{1} \zeta_{2}}+\eta\left(\zeta_{2}\right) \zeta_{1}-g\left(\zeta_{1}, \zeta_{2}\right) \xi$ for all $\zeta_{1}$ and $\zeta_{2}$ on $M^{n}$, where $\eta$ is a 1 -form associated with the unit vector field $\xi$ such that $g(\xi, \xi)=\eta(\xi)=1$ and $g\left(\zeta_{1}, \xi\right)=\eta\left(\zeta_{1}\right)$. Later, the properties of the semi-symmetric metric connection $\tilde{\nabla}$ have been explored by several researchers. One of these properties is the curvature invariant respecting to the semi-symmetric metric connection $\tilde{\nabla}$ and the Levi-Civita connection $\nabla$. For example, the conformal curvature tensors corresponding to the semi-symmetric connection (Yano's sense) and the Levi-Civita connection coincide. Similar results for different curvature tensors have been established by many geometers. A connection $\tilde{\nabla}$ is said to be semi-symmetric metric $\xi$-connection if and only if $\tilde{\nabla} \xi=0$. Afterwards, the properties of semi-symmetric metric $\tilde{\xi}$-connection have been studied in [21-24].

In an $\left(M^{n}, \tilde{\nabla}, g\right)$, we have [21]

$$
\begin{equation*}
\nabla_{\zeta_{1}} \xi=-\zeta_{1}+\eta\left(\zeta_{1}\right) \xi, \quad g(\xi, \xi)=1, \text { and } \eta\left(\zeta_{1}\right)=g\left(\zeta_{1}, \xi\right) \tag{4}
\end{equation*}
$$

for any $\zeta_{1}$ on $M^{n}$. Next, we have [21]

$$
\begin{gather*}
\left(\nabla_{\zeta_{1}} \eta\right) \zeta_{2}=-g\left(\zeta_{1}, \zeta_{2}\right)+\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right),  \tag{5}\\
K\left(\zeta_{1}, \zeta_{2}\right) \xi=\eta\left(\zeta_{1}\right) \zeta_{2}-\eta\left(\zeta_{2}\right) \zeta_{1},  \tag{6}\\
K\left(\zeta_{1}, \xi\right) \zeta_{2}=g\left(\zeta_{1}, \zeta_{2}\right) \xi-\eta\left(\zeta_{2}\right) \zeta_{1},  \tag{7}\\
S\left(\zeta_{1}, \xi\right)=-(n-1) \eta\left(\zeta_{1}\right) \Longleftrightarrow Q \xi=-(n-1) \xi,  \tag{8}\\
\left(£_{\xi} g\right)\left(\zeta_{1}, \zeta_{2}\right)=2\left\{-g\left(\zeta_{1}, \zeta_{2}\right)+\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)\right\}, \tag{9}
\end{gather*}
$$

for all $\zeta_{1}, \zeta_{2}$ on $M^{n}$. Here, $K$ and $Q$ represent the curvature tensor and the Ricci operator of $M^{n}$, respectively.

Definition 1. An $M^{n}$ is said to be quasi-Einstein if its $S(\neq 0)$ satisfies

$$
S\left(\zeta_{1}, \zeta_{2}\right)=\mathfrak{l} g\left(\zeta_{1}, \zeta_{2}\right)+\mathfrak{m} \eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)
$$

where $\mathfrak{m}$ and $\mathfrak{l}$ are smooth functions on $M^{n}$. If $\mathfrak{m}=0$, then the manifold is called an Einstein manifold.

Definition 2. A partial differential equation $\Delta u=v$ on a complete $M^{n}$ is called a Poisson equation for some smooth functions $u$ and $v$.

Remark 1 ([21,22]). An $\left(M^{3}, \tilde{\nabla}, g\right)$ is a quasi-Einstein manifold of the form

$$
\begin{equation*}
S\left(\zeta_{1}, \zeta_{2}\right)=\left(1+\frac{R}{2}\right) g\left(\zeta_{1}, \zeta_{2}\right)-\left(3+\frac{R}{2}\right) \eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right) \tag{10}
\end{equation*}
$$

Remark $2([21,22])$. In an $\left(M^{3}, \tilde{\nabla}, g\right)$, we have

$$
\begin{align*}
\xi(R) & =2(R+6)  \tag{11}\\
\eta\left(\nabla_{\xi} \mathcal{D} R\right) & =4(R+6) \tag{12}
\end{align*}
$$

where $\mathcal{D}$ is the gradient operator of $g$. From (11), it is noticed that $R$ of $M^{3}$ is constant if and only if $R=-6$.

## 3. $\rho$-ES on $\left(M^{3}, \tilde{\nabla}, g\right)$

First, we prove the following theorem.
Theorem 1. If $\left(M^{3}, \tilde{\nabla}, g\right)$ admits a $\rho-E S(g, E, \sigma, \rho)$, then its scalar curvature $R$ satisfies the Poisson equation $\Delta R=\frac{4(2-\sigma-6 \rho)}{\rho}$, provided $\rho \neq 0$.

Proof. Let the metric of an $\left(M^{3}, \tilde{\nabla}, g\right)$ be a $\rho$-ES $(g, E, \sigma, \rho)$, then in view of (10), (2) leads to

$$
\begin{align*}
\left(£_{E} g\right)\left(\zeta_{1}, \zeta_{2}\right)= & -2\left\{1+\sigma+\left(\frac{1}{2}-\rho\right) R\right\} g\left(\zeta_{1}, \zeta_{2}\right)  \tag{13}\\
& +(R+6) \eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)
\end{align*}
$$

for any vector fields $\zeta_{1}, \zeta_{2}$ on $M^{3}$.
Taking covariant derivative of (13) respecting to $\zeta_{3}$, we find

$$
\begin{align*}
\left(\nabla_{\zeta_{3}} £_{E} g\right)\left(\zeta_{1}, \zeta_{2}\right)= & \left(\zeta_{3} R\right)\left\{(2 \rho-1) g\left(\zeta_{1}, \zeta_{2}\right)+\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)\right\}  \tag{14}\\
& -(R+6)\left\{g\left(\zeta_{1}, \zeta_{3}\right) \eta\left(\zeta_{2}\right)+g\left(\zeta_{2}, \zeta_{3}\right) \eta\left(\zeta_{1}\right)-2 \eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right) \eta\left(\zeta_{3}\right)\right\}
\end{align*}
$$

As $g$ is parallel with respect to $\nabla$, then the formula [25]

$$
\left(£_{E} \nabla_{\zeta_{1}} g-\nabla_{\zeta_{1}} £_{E} g-\nabla_{\left[E, \zeta_{1}\right]} g\right)\left(\zeta_{2}, \zeta_{3}\right)=-g\left(\left(£_{E} \nabla\right)\left(\zeta_{1}, \zeta_{2}\right), \zeta_{3}\right)-g\left(\left(£_{E} \nabla\right)\left(\zeta_{1}, \zeta_{3}\right), \zeta_{2}\right)
$$

turns to

$$
\left(\nabla_{\zeta_{1}} £_{E} g\right)\left(\zeta_{2}, \zeta_{3}\right)=g\left(\left(£_{E} \nabla\right)\left(\zeta_{1}, \zeta_{2}\right), \zeta_{3}\right)+g\left(\left(£_{E} \nabla\right)\left(\zeta_{1}, \zeta_{3}\right), \zeta_{2}\right)
$$

Since $£_{E} \nabla$ is symmetric, therefore we have

$$
2 g\left(\left(£_{E} \nabla\right)\left(\zeta_{1}, \zeta_{2}\right), \zeta_{3}\right)=\left(\nabla_{\zeta_{1}} £_{E} g\right)\left(\zeta_{2}, \zeta_{3}\right)+\left(\nabla_{\zeta_{2}} £_{E} g\right)\left(\zeta_{1}, \zeta_{3}\right)-\left(\nabla_{\zeta_{3}} £_{E} g\right)\left(\zeta_{1}, \zeta_{2}\right),
$$

which in view of (14) gives

$$
\begin{aligned}
2 g\left(\left(£_{E} \nabla\right)\left(\zeta_{1}, \zeta_{2}\right), \zeta_{3}\right)= & \left(\zeta_{1} R\right)\left\{(2 \rho-1) g\left(\zeta_{2}, \zeta_{3}\right)+\eta\left(\zeta_{2}\right) \eta\left(\zeta_{3}\right)\right\} \\
& +\left(\zeta_{2} R\right)\left\{(2 \rho-1) g\left(\zeta_{1}, \zeta_{3}\right)+\eta\left(\zeta_{1}\right) \eta\left(\zeta_{3}\right)\right\} \\
& -\left(\zeta_{3} R\right)\left\{(2 \rho-1) g\left(\zeta_{1}, \zeta_{2}\right)+\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)\right\} \\
& -2(R+6)\left\{g\left(\zeta_{1}, \zeta_{2}\right) \eta\left(\zeta_{3}\right)-\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right) \eta\left(\zeta_{3}\right)\right\},
\end{aligned}
$$

from which it follows that

$$
\begin{align*}
2\left(£_{E} \nabla\right)\left(\zeta_{1}, \zeta_{2}\right)= & \left(\zeta_{1} R\right)\left\{(2 \rho-1) \zeta_{2}+\eta\left(\zeta_{2}\right) \xi\right\}  \tag{15}\\
& +\left(\zeta_{2} R\right)\left\{(2 \rho-1) \zeta_{1}+\eta\left(\zeta_{1}\right) \xi\right\} \\
& -(\mathcal{D} R)\left\{(2 \rho-1) g\left(\zeta_{1}, \zeta_{2}\right)+\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)\right\} \\
& -2(R+6)\left\{g\left(\zeta_{1}, \zeta_{2}\right) \xi-\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right) \xi\right\}
\end{align*}
$$

Replacing $\zeta_{2}$ by $\xi$ and $\zeta_{1}$ by $\zeta_{2}$ in (15), we have

$$
\begin{align*}
\left(£_{E} \nabla\right)\left(\zeta_{2}, \xi\right)= & \rho g\left(\mathcal{D} R, \zeta_{2}\right) \xi-\rho(\mathcal{D} R) \eta\left(\zeta_{2}\right)  \tag{16}\\
& +(R+6)\left\{(2 \rho-1) \zeta_{2}+\eta\left(\zeta_{2}\right) \xi\right\} .
\end{align*}
$$

The covariant differentiation of (16) respecting to $\zeta_{1}$ yields

$$
\begin{align*}
\left(\nabla_{\zeta_{1}} £_{E} \nabla\right)\left(\zeta_{2}, \xi\right)= & 2\left(\zeta_{1} R\right)\left\{(2 \rho-1) \zeta_{2}+\eta\left(\zeta_{2}\right) \xi\right\} \\
& +\left(\zeta_{2} R\right)\left\{(\rho-1) \zeta_{1}+\eta\left(\zeta_{1}\right) \xi\right\} \\
& -(\mathcal{D} R)\left\{(\rho-1) g\left(\zeta_{1}, \zeta_{2}\right)+\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)\right\}  \tag{17}\\
& -3(R+6)\left\{g\left(\zeta_{1}, \zeta_{2}\right) \xi-\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right) \xi\right\} \\
& -(R+6)\left\{(2 \rho-1) \eta\left(\zeta_{1}\right) \zeta_{2}+\eta\left(\zeta_{2}\right) \zeta_{1}\right\} \\
& +\rho g\left(\nabla_{\zeta_{1}} \mathcal{D} R, \zeta_{2}\right) \xi-\rho\left(\nabla_{\zeta_{1}} \mathcal{D} R\right) \eta\left(\zeta_{2}\right),
\end{align*}
$$

where (4), (5) and (16) being used.
Again from [25], we have

$$
\begin{equation*}
\left(£_{E} K\right)\left(\zeta_{1}, \zeta_{2}\right) \zeta_{3}=\left(\nabla_{\zeta_{1}} £_{E} \nabla\right)\left(\zeta_{2}, \zeta_{3}\right)-\left(\nabla_{\zeta_{2}} £_{E} \nabla\right)\left(\zeta_{1}, \zeta_{3}\right), \tag{18}
\end{equation*}
$$

which by putting $\zeta_{3}=\xi$ and using (17) becomes

$$
\begin{align*}
\left(£_{E} K\right)\left(\zeta_{1}, \zeta_{2}\right) \xi= & g\left(\mathcal{D} R, \zeta_{1}\right)\left\{(3 \rho-1) \zeta_{2}+\eta\left(\zeta_{2}\right) \xi\right\}  \tag{19}\\
& -g\left(\mathcal{D} R, \zeta_{2}\right)\left\{(3 \rho-1) \zeta_{1}+\eta\left(\zeta_{1}\right) \xi\right\} \\
& +2(R+6)(\rho-1)\left\{\eta\left(\zeta_{2}\right) \zeta_{1}-\eta\left(\zeta_{1}\right) \zeta_{2}\right\} \\
& +\rho g\left(\nabla_{\zeta_{1}} \mathcal{D} R, \zeta_{2}\right) \xi-\rho g\left(\nabla_{\zeta_{2}} \mathcal{D} R, \zeta_{1}\right) \xi \\
& -\rho\left(\nabla_{\zeta_{1}} \mathcal{D} R\right) \eta\left(\zeta_{2}\right)+\rho\left(\nabla_{\zeta_{2}} \mathcal{D} R\right) \eta\left(\zeta_{1}\right) .
\end{align*}
$$

Contracting (19) respecting to $\zeta_{1}$ then using (4) and (11) we lead to

$$
\begin{align*}
\left(£_{E} S\right)\left(\zeta_{2}, \xi\right)= & (1-6 \rho) \zeta_{2}(R)+2(R+6)(2 \rho-1) \eta\left(\zeta_{2}\right)  \tag{20}\\
& +\rho g\left(\nabla_{\xi} \mathcal{D} R, \zeta_{2}\right) \xi-\rho(\Delta R) \eta\left(\zeta_{2}\right) .
\end{align*}
$$

By putting $\zeta_{2}=\xi$ in (20) then using (4), (11) and (12), we find

$$
\begin{equation*}
\left(£_{E} S\right)(\xi, \xi)=-4 \rho(R+6)-\rho(\Delta R) . \tag{21}
\end{equation*}
$$

The Lie derivative of (8) respecting to $E$ leads to

$$
\begin{equation*}
\left(£_{E} S\right)(\xi, \xi)=4 \eta\left(£_{E} \xi\right) . \tag{22}
\end{equation*}
$$

Putting $\zeta_{1}=\zeta_{2}=\xi$ in (13) infers

$$
\begin{equation*}
\left(£_{E} g\right)(\xi, \xi)=-2 \sigma+2 \rho R+4 \tag{23}
\end{equation*}
$$

The Lie derivative of $g(\xi, \xi)=1$ gives

$$
\begin{equation*}
\left(£_{E} g\right)(\xi, \xi)=-2 \eta\left(£_{E} \xi\right) . \tag{24}
\end{equation*}
$$

Now combining (21)-(25) we deduce

$$
\begin{equation*}
\Delta R=\frac{4(2-\sigma-6 \rho)}{\rho}, \text { provided } \rho \neq 0 \tag{25}
\end{equation*}
$$

This completes the proof.
It is well-known that the $\rho$-ES Equation (2) on $M^{n}$ with the soliton constant $\rho=\frac{1}{2}, \frac{1}{n}, \frac{1}{2(n-1)}$ reduces to the Einstein soliton, traceless Ricci soliton, Schouten soliton, respectively. It is also known that a smooth function $\mathfrak{f}$ on an $M^{n}$ is called harmonic, subharmonic or superharmonic if $\Delta \mathfrak{f}=0, \geq 0$ or $\leq 0$, respectively. These facts together with Theorem 1 state the following:

Corollary 1. Let $\left(M^{3}, \tilde{\nabla}, g\right)$ admit a $\rho-E S$, then we have

| Value of $\rho$ | Solitons | Poisson equation | Condition for $R$ to be subharmonic <br> and superharmonic |
| :--- | :--- | :--- | :--- |
| $\rho=\frac{1}{2}$ | Einstein soliton | $\Delta R=-8(\sigma+1)$ | (i) R is subharmonic if $\sigma \leq-1$, <br> (ii) $R$ is superharmonic if $\sigma \geq-1$, |
| $\rho=\frac{1}{3}$ | traceless Ricci soliton | $\Delta R=-12 \sigma$ | (i) $R$ is subharmonic if $\sigma \leq 0$, <br> (ii) $R$ is superharmonic if $\sigma \geq 0$, |
| $\rho=\frac{1}{4}$ | Schouten soliton | $\Delta R=16\left(\frac{1}{2}-\sigma\right)$ | (i) $R$ is subharmonic if $\sigma \leq \frac{1}{2}$, <br> (ii) $R$ is superharmonic if $\sigma \geq \frac{1}{2}$. |

Remark 3. The $\rho$-ES on an $M^{n}$ with $\rho=0$ reduces to the RS. The properties of RS on $\left(M^{3}, \tilde{\nabla}, g\right)$ have been explored by Chaubey and De [22]. Thus, we can say that the Theorem 1 generalizes the study of Einstein soliton, traceless RS and the Schouten soliton on $\left(M^{3}, \tilde{\nabla}, g\right)$.

It is well-known that the Poisson equation $\Delta u=v$ with $v=0$ becomes a Laplace equation. Suppose that an $\left(M^{3}, \tilde{\nabla}, g\right)$ does not admit RS. Then, Theorem 1 and above discussion state:

Corollary 2. If $\left(M^{3}, \tilde{\nabla}, g\right)$ admits a $\rho-E S$, which is not a RS $(\rho \neq 0)$, then $R$ of $M^{3}$ satisfies Laplace equation if and only if $\sigma=2(1-3 \rho)$.

Let $\left(M^{3}, \tilde{\nabla}, g\right)$ admit a $\rho$-ES. If $R$ of $M^{3}$ satisfies the Laplace equation, then $\sigma=2(1-3 \rho)$. The $\rho$-ES under consideration to be steady, shrinking or expanding if $\rho$ is equal to, less than or greater than $\frac{1}{3}$. Thus, we write our corollary as

Corollary 3. Let the metric of an $\left(M^{3}, \tilde{\nabla}, g\right)$ be $\rho$-ES, which is not a RS $(\rho \neq 0)$. If $R$ of $M^{3}$ satisfies the Laplace equation, then the $\rho$-ES is steady, shrinking or expanding if $\rho=\frac{1}{3}, \rho<\frac{1}{3}$ or $\rho>\frac{1}{3}$, respectively.

## 4. Gradient $\rho$-ES on $\left(M^{3}, \tilde{\nabla}, g\right)$

Theorem 2. Let $\left(M^{3}, \tilde{\nabla}, g\right)$ admit a gradient $\rho-E S$. Then, either $M^{3}$ is Einstein or the gradient $\rho$-ES is steady type gradient traceless RS.

Proof. Let the metric of an $\left(M^{3}, \tilde{\nabla}, g\right)$ be a gradient $\rho$-ES. Then, (3) can be written as

$$
\begin{equation*}
\nabla_{\zeta_{1}} \mathcal{D} \mathcal{F}+Q \zeta_{1}+(\sigma-\rho R) \zeta_{1}=0 \tag{26}
\end{equation*}
$$

for all $\zeta_{1}$ on $M^{3}$.
The covariant differentiation of (26) with respect to $\zeta_{2}$ leads to

$$
\begin{equation*}
\nabla_{\zeta_{2}} \nabla_{\zeta_{1}} \mathcal{D F}=-\left(\nabla_{\zeta_{2}} Q\right) \zeta_{1}-Q\left(\nabla_{\zeta_{2}} \zeta_{1}\right)-(\sigma-\rho R) \nabla_{\zeta_{2}} \zeta_{1}+\rho \zeta_{2}(R) \zeta_{1} \tag{27}
\end{equation*}
$$

Interchanging $\zeta_{1}$ and $\zeta_{2}$ in (27) leads to

$$
\begin{equation*}
\nabla_{\zeta_{1}} \nabla_{\zeta_{2}} \mathcal{D} \mathcal{F}=-\left(\nabla_{\zeta_{1}} Q\right) \zeta_{2}-Q\left(\nabla_{\zeta_{1}} \zeta_{2}\right)-(\sigma-\rho R) \nabla_{\zeta_{1}} \zeta_{2}+\rho \zeta_{1}(R) \zeta_{2} \tag{28}
\end{equation*}
$$

By plugging of (26)-(28), we find

$$
K\left(\zeta_{1}, \zeta_{2}\right) \mathcal{D \mathcal { F }}=-\left(\nabla_{\zeta_{1}} Q\right) \zeta_{2}+\left(\nabla_{\zeta_{2}} Q\right) \zeta_{1}+\rho\left\{\zeta_{1}(R) \zeta_{2}-\zeta_{2}(R) \zeta_{1}\right\}
$$

Contracting the forgoing equation along $\zeta_{1}$, we obtain

$$
\begin{equation*}
S\left(\zeta_{2}, \mathcal{D} \mathcal{F}\right)=\frac{(1-4 \rho)}{2} \zeta_{2}(R) \tag{29}
\end{equation*}
$$

In account of (10), we have

$$
\begin{equation*}
S\left(\zeta_{2}, \mathcal{D} \mathcal{F}\right)=\left(1+\frac{R}{2}\right) \zeta_{2}(\mathcal{F})-\left(3+\frac{R}{2}\right) \eta\left(\zeta_{2}\right) \xi(\mathcal{F}) \tag{30}
\end{equation*}
$$

Thus, from (29) and (30), it follows that

$$
\begin{equation*}
(1-4 \rho) \zeta_{2}(R)=(R+2) \zeta_{2}(\mathcal{F})-(R+6) \eta\left(\zeta_{2}\right) \xi(\mathcal{F}) \tag{31}
\end{equation*}
$$

By putting $\zeta_{2}=\xi$ in (31), then using (4) and (11), we find

$$
\begin{equation*}
\xi(\mathcal{F})=-\frac{1}{2}(1-4 \rho)(R+6) \tag{32}
\end{equation*}
$$

By using (32) and (31) turns to

$$
\begin{equation*}
(1-4 \rho) \zeta_{2}(R)=(R+2) \zeta_{2}(\mathcal{F})+\frac{1}{2}(R+6)^{2}(1-4 \rho) \eta\left(\zeta_{2}\right) \tag{33}
\end{equation*}
$$

The covariant differentiation of (33) along $\zeta_{1}$ leads to

$$
\begin{align*}
(1-4 \rho) g\left(\nabla_{\zeta_{1}} \mathcal{D} R, \zeta_{2}\right)= & \zeta_{1}(R) \zeta_{2}(\mathcal{F})+(R+2) g\left(\nabla_{\zeta_{1}} \mathcal{D} \mathcal{F}, \zeta_{2}\right)  \tag{34}\\
& +(R+6)(1-4 \rho) \zeta_{1}(R) \eta\left(\zeta_{2}\right) \\
& +\frac{1}{2}(R+6)^{2}(1-4 \rho)\left\{\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)-g\left(\zeta_{1}, \zeta_{2}\right)\right\} .
\end{align*}
$$

Interchanging $\zeta_{1}$ and $\zeta_{2}$ in (34), we have

$$
\begin{align*}
(1-4 \rho) g\left(\nabla_{\zeta_{2}} \mathcal{D} R, \zeta_{1}\right)= & \zeta_{2}(R) \zeta_{1}(\mathcal{F})+(R+2) g\left(\nabla_{\zeta_{2}} \mathcal{D} \mathcal{F}, \zeta_{1}\right)  \tag{35}\\
& +(R+6)(1-4 \rho) \zeta_{2}(R) \eta\left(\zeta_{1}\right) \\
& +\frac{1}{2}(R+6)^{2}(1-4 \rho)\left\{\eta\left(\zeta_{1}\right) \eta\left(\zeta_{2}\right)-g\left(\zeta_{1}, \zeta_{2}\right)\right\}
\end{align*}
$$

Equating the left hand sides of last two equations gives

$$
\begin{gathered}
\zeta_{1}(R) \zeta_{2}(\mathcal{F})+(R+6)(1-4 \rho) \zeta_{1}(R) \eta\left(\zeta_{2}\right) \\
-\zeta_{2}(R) \zeta_{1}(\mathcal{F})-(R+6)(1-4 \rho) \zeta_{2}(R) \eta\left(\zeta_{1}\right)=0
\end{gathered}
$$

which by replacing $\zeta_{2}=\xi$ then using (4), (11) and (32) takes the form

$$
(R+6)\left\{(1-4 \rho) \zeta_{1}(R)-4 \zeta_{1}(\mathcal{F})-4(R+6)(1-4 \rho) \eta\left(\zeta_{1}\right)\right\}=0
$$

Thus, we have either $R=-6$, or $(1-4 \rho) \zeta_{1}(R)=4 \zeta_{1}(\mathcal{F})+4(R+6)(1-4 \rho) \eta\left(\zeta_{1}\right)$. If we firstly suppose that $R \neq-6$ and $(1-4 \rho) \zeta_{1}(R)=4 \zeta_{1}(\mathcal{F})+4(R+6)(1-4 \rho) \eta\left(\zeta_{1}\right)$, which by virtue of (33) turns to

$$
\begin{equation*}
(R-2)\left\{2 \zeta_{1}(\mathcal{F})+(R+6)(1-4 \rho) \eta\left(\zeta_{1}\right)\right\}=0, \tag{36}
\end{equation*}
$$

which refers that either $R=2$ or $\zeta_{1}(\mathcal{F})=-\frac{1}{2}(R+6)(1-4 \rho) \eta\left(\zeta_{1}\right)$. From (11), it is obvious that if $R$ is constant, then its value must be -6 , which shows that $R=2$ is inadmissible. Thus, we have $\zeta_{1}(\mathcal{F})=-\frac{1}{2}(R+6)(1-4 \rho) \eta\left(\zeta_{1}\right)$, which is equivalent to

$$
\begin{equation*}
\mathcal{D} \mathcal{F}=-\frac{1}{2}(R+6)(1-4 \rho) \xi=\xi(\mathcal{F}) \xi \tag{37}
\end{equation*}
$$

Thus, the gradient of $\mathcal{F}$ is pointwise collinear with $\xi$. Now, taking the covariant derivative of (37) with respect to $\zeta_{1}$ and using (4), we have

$$
\begin{equation*}
\nabla_{\zeta_{1}} \mathcal{D} \mathcal{F}=\zeta_{1}(\xi(\mathcal{F})) \xi-\xi(\mathcal{F})\left(\zeta_{1}-\eta\left(\zeta_{1}\right) \xi\right) \tag{38}
\end{equation*}
$$

Therefore, from (26) and (38), we obtain

$$
\begin{equation*}
Q \zeta_{1}+(\sigma-\rho R) \zeta_{1}=-\zeta_{1}(\xi(\mathcal{F})) \xi+\xi(\mathcal{F})\left(\zeta_{1}-\eta\left(\zeta_{1}\right) \xi\right) \tag{39}
\end{equation*}
$$

Now, by replacing $\zeta_{1}$ by $\xi$ in (39) then using (8), (11) and (32) we lead to

$$
\begin{equation*}
\sigma=(1-3 \rho)(R+8) \tag{40}
\end{equation*}
$$

Let us suppose that $\rho=\frac{1}{3}$, that is, the gradient $\rho$-ES on an $M^{3}$ is gradient traceless RS. This fact together with Equation (40) leads to $\sigma=0$. Thus, the gradient traceless RS is steady. This completes the proof.

Theorem 3. Let an $\left(M^{3}, \tilde{\nabla}, g\right)$ be a non-gradient traceless RS. Then, the gradient $\rho$-ES is trivial soliton with constant $\sigma=2(1-3 \rho)$. Also, the $\rho-E S$ is shrinking and expanding according to $\rho>\frac{1}{3}$ and $\rho<\frac{1}{3}$.

Proof. Now, we suppose that $\rho \neq \frac{1}{3}$. Thus, (40) leads to

$$
\begin{equation*}
R=\frac{\sigma}{1-3 \rho}-8 \tag{41}
\end{equation*}
$$

which informs that $R$ is constant and hence (11) infers that $R=-6$. This contradicts our hypothesis $R \neq-6$.

Secondly, we consider that $R=-6$ and $(1-4 \rho) \zeta_{1}(R) \neq 4 \zeta_{1}(\mathcal{F})+4(R+6)(1-$ $4 \rho) \eta\left(\zeta_{1}\right)$. For $R=-6$, (33) informs that $\mathcal{F} \in \mathcal{R}$ and hence the GRBS on the manifold is trivial. Moreover, the Riemannian 3-manifold under assumption is an Einstein manifold with $\sigma=2(1-3 \rho)$. This completes the proof.

Let us suppose that an $\left(M^{3}, \tilde{\nabla}, g\right)$ admits a proper gradient $\rho$-ES. Then, the $\rho$-ES reduces to the gradient traceless RS and $\rho=\frac{1}{3}, \sigma=0$. Using these facts in (26) and then contracting the foregoing equation over $\zeta_{1}$ gives $\Delta \mathcal{F}=0$.

A smooth function $\mathfrak{h}$ on an $M^{n}$ is called harmonic if $\Delta \mathfrak{h}=0$.
The above discussions state the following:
Corollary 4. Let a complete $\left(M^{3}, \tilde{\nabla}, g\right)$ admit a proper gradient $\rho-E S$. Then the gradient function of the gradient $\rho$-ES is harmonic.

Contracting (38) over $\zeta_{1}$, we find

$$
\Delta \mathcal{F}=\xi(\xi(\mathcal{F}))-2 \xi(\mathcal{F})
$$

Again, considering $\sigma=0, \rho=\frac{1}{3}$ and then contracting (26) over $\zeta_{1}$, we conclude that

$$
\Delta \mathcal{F}=0
$$

The last two equations show that $\xi(\xi(\mathcal{F}))-2 \xi(\mathcal{F})=0$. Let $\xi=\frac{\partial}{\partial t}$. Thus, we notice that the potential function $\mathcal{F}$ satisfies the PDE

$$
\frac{\partial^{2} \mathcal{F}}{\partial t^{2}}-2 \frac{\partial \mathcal{F}}{\partial t}=0
$$

It is obvious that $\mathcal{F}=A e^{2 t}+B$ for smooth functions $A$ and $B$, which are independent of $t$, is the solution of the above PDE. Now, we list our results in the following:

Corollary 5. Let the metric of a complete $\left(M^{3}, \tilde{\nabla}, g\right)$ admit a proper gradient $\rho$-ES. Then, the potential function $\mathcal{F}$ of such soliton satisfies the $P D E \frac{\partial^{2} \mathcal{F}}{\partial t^{2}}-2 \frac{\partial \mathcal{F}}{\partial t}=0$, and it can be evaluated by $\mathcal{F}=A e^{2 t}+B$.

## 5. Example

We consider the manifold $M^{3}=\left\{\left(w_{1}, w_{2}, w_{3}\right) \in \mathcal{R}^{3}\right\}$, where $\left(w_{1}, w_{2}, w_{3}\right)$ are the usual coordinates in $\mathbb{R}^{3}$. Let $u_{1}, u_{2}$ and $u_{3}$ be the vector fields on $M^{3}$ given by

$$
u_{1}=e^{\mathfrak{b} w_{3}+w_{1}} \frac{\partial}{\partial w_{1}}, \quad u_{2}=e^{\mathfrak{b} w_{3}+w_{2}} \frac{\partial}{\partial w_{2}}, \quad u_{3}=\frac{1}{\mathfrak{b}} \frac{\partial}{\partial w_{3}}=\xi
$$

where $\mathfrak{b}(\neq 0) \in \mathcal{R}$. Then, $\left\{u_{1}, u_{2}, u_{3}\right\}$ forms a basis in the module of the vector fields of $M^{3}$.
Let the Riemannian metric $g$ be defined by

$$
g\left(u_{p}, u_{q}\right)= \begin{cases}1, & 1 \leq p=q \leq 3 \\ 0, & \text { otherwise }\end{cases}
$$

Hence, $M^{3}$ is a Riemannian manifold of dimension 3. Let the 1-form $\eta$ on $M^{3}$ be defined by $\eta\left(\zeta_{1}\right)=g\left(\zeta_{1}, u_{3}\right)=g\left(\zeta_{1}, \xi\right)$ for all $\zeta_{1}$ on $M^{3}$. Now, by direct computations, we obtain

$$
\left[u_{1}, u_{2}\right]=0, \quad\left[u_{1}, u_{3}\right]=-u_{1}, \quad\left[u_{2}, u_{3}\right]=-u_{2} .
$$

By using Koszul's formula, we obtain

$$
\nabla_{u_{p}} u_{q}=\left\{\begin{array}{l}
-u_{p}, \quad p=1,2, q=3 \\
u_{3}, \quad 1 \leq p=q \leq 2 \\
0, \quad \text { otherwise }
\end{array}\right.
$$

Now we suppose that $\zeta_{1}=\zeta_{1}^{1} u_{1}+\zeta_{1}^{2} u_{2}+\zeta_{1}^{3} u_{3}$, then for $\xi=u_{3}$ it follows that $\nabla_{\zeta_{1}} \xi=-\zeta_{1}+\eta\left(\zeta_{1}\right) \xi$. It can be easily seen that $\tilde{\nabla}$ defined on $M^{3}$ satisfies the conditions

$$
\tilde{T}\left(\zeta_{1}, \zeta_{2}\right)=-\eta\left(\zeta_{1}\right) \zeta_{2}+\eta\left(\zeta_{2}\right) \zeta_{1}, \tilde{\nabla} g=0, \text { and } \tilde{\nabla} \xi=0
$$

for arbitrary vector fields $\zeta_{1}$ and $\zeta_{2}$ on $M^{3}$, where $\tilde{T}$ indicates the torsion tensor of $\tilde{\nabla}$. Thus, we can say that $\tilde{\nabla}$ is a semi-symmetric metric $\xi$-connection on $M^{3}$.

The non-zero constituents of $K$ are obtained as follows:

$$
\begin{gathered}
K\left(u_{1}, u_{3}\right) u_{1}=u_{3}, \quad K\left(u_{1}, u_{2}\right) u_{1}=u_{2}, \quad K\left(u_{2}, u_{3}\right) u_{2}=u_{3}, \\
K\left(u_{1}, u_{2}\right) u_{2}=K\left(u_{1}, u_{3}\right) u_{3}=-u_{1}, \quad K\left(u_{2}, u_{3}\right) u_{3}=-u_{2} .
\end{gathered}
$$

By using above components of the curvature tensor $K$ we obtain

$$
S\left(u_{p}, u_{q}\right)=-2, \quad 1 \leq p=q \leq 3
$$

from which we obtain $R=-6$.
Now, by taking $\mathcal{D} \mathcal{F}=\left(u_{1} \mathcal{F}\right) u_{1}+\left(u_{2} \mathcal{F}\right) u_{2}+\left(u_{3} \mathcal{F}\right) u_{3}$, we have

$$
\begin{gathered}
\nabla_{u_{1}} \mathcal{D F}=\left(u_{1}\left(u_{1} \mathcal{F}\right)-u_{3} \mathcal{F}\right) u_{1}+\left(u_{1}\left(u_{2} \mathcal{F}\right)\right) u_{2}+\left(u_{1}\left(u_{3} \mathcal{F}\right)+u_{1} \mathcal{F}\right) u_{3}, \\
\nabla_{\mathcal{E}_{2}} \mathcal{D} \mathcal{F}=\left(u_{2}\left(u_{1} \mathcal{F}\right)\right) u_{1}+\left(u_{2}\left(u_{2} \mathcal{F}\right)-u_{3} \mathcal{F}\right) u_{2}+\left(u_{2}\left(u_{3} \mathcal{F}\right)+\mathcal{F}_{2} \mathcal{F}\right) \mathcal{F}_{3}, \\
\nabla_{\mathcal{E}_{3}} \mathcal{D} \mathcal{F}=\left(u_{3}\left(u_{1} \mathcal{F}\right)\right) u_{1}+\left(u_{3}\left(u_{2} \mathcal{F}\right)\right) u_{2}+\left(u_{3}\left(u_{3} \mathcal{F}\right)\right) u_{3} .
\end{gathered}
$$

Thus, by virtue of (26), we obtain

$$
\left\{\begin{array}{l}
u_{1}\left(u_{1} \mathcal{F}\right)-u_{3} \mathcal{F}=2-6 \rho-\sigma,  \tag{42}\\
u_{2}\left(u_{2} \mathcal{F}\right)-u_{3} \mathcal{F}=2-6 \rho-\sigma, \\
u_{3}\left(u_{3} \mathcal{F}\right)=2-6 \rho-\sigma, \\
u_{1}\left(u_{2} \mathcal{F}\right)=0, \\
u_{2}\left(u_{1} \mathcal{F}\right)=0, \\
u_{2}\left(u_{3} \mathcal{F}\right)+u_{2} \mathcal{F}=0
\end{array}\right.
$$

Thus, the relations in (42) are, respectively, amounting to

$$
\begin{gathered}
e^{2\left(\mathfrak{b} w_{3}+w_{1}\right)}\left[\frac{\partial^{2} \mathcal{F}}{\partial w_{1}^{2}}+\frac{\partial \mathcal{F}}{\partial w_{1}}\right]-\frac{1}{\mathfrak{b}} \frac{\partial \mathcal{F}}{\partial w_{3}}=2-6 \rho-\sigma, \\
e^{2\left(\mathfrak{b} w_{3}+w_{1}\right)}\left[\frac{\partial^{2} \mathcal{F}}{\partial w_{2}^{2}}+\frac{\partial \mathcal{F}}{\partial w_{2}}\right]-\frac{1}{\mathfrak{b}} \frac{\partial \mathcal{F}}{\partial w_{3}}=2-6 \rho-\sigma, \\
\frac{1}{\mathfrak{b}^{2}} \frac{\partial^{2} \mathcal{F}}{\partial w_{3}^{2}}=2-6 \rho-\sigma \\
\frac{\partial^{2} \mathcal{F}}{\partial w_{1} \partial w_{2}}=0 \\
\frac{\partial^{2} \mathcal{F}}{\partial w_{2} \partial w_{1}}=0 \\
\frac{1}{\mathfrak{b}}\left[\frac{\partial^{2} \mathcal{F}}{\partial w_{2} \partial w_{3}}+\frac{\partial \mathcal{F}}{\partial w_{2}}\right]=0
\end{gathered}
$$

From the above relations, it is noticed that $\mathcal{F} \in \mathcal{R}$ for $\sigma=2-6 \rho$. Hence, the Equation (26) is satisfied. Thus, $g$ is a gradient $\rho$-ES with the soliton vector field $E=\mathcal{D} \mathcal{F}$, where $\mathcal{F} \in \mathcal{R}$ and $\sigma=2-6 \rho$. For $\rho=\frac{1}{3}$, we obtain $\sigma=0$, i.e., the gradient $\rho$-ES is trivial with constant $\sigma=2-6 \rho$. Thus, Theorem 2 is verified.

## 6. Results and Discussion

It is well known that the $\rho$-Einstein soliton Equation (2) with $\rho=0$ becomes the Ricci soliton equation, which has been studied in [22]. Thus, we can say that the $\rho$-Einstein soliton is a natural generalization of Ricci soliton. In this manuscript, we have explored the properties of $\rho$-Einstein solitons in Riemannian geometry, which generalizes the results of [22].

## 7. Conclusions

To prove the curvatures invariant, Chauey et al. [23] defined the notion of semisymmetric metric $P$-connection in Riemannian setting, which is a particular case of Riemannian concircular structure manifold [26]. This topic has great applications in differential equations. We proved that the scalar curvature of Riemannian 3-manifolds endowed with a semi-symmetric metric $\xi$-connection and Ricci-Bourguignon soliton satisfies the Poisson and Laplace equations. It is well known that the Poisson and Laplace equations play a crucial role in the development of engineering, physics, mathematics, etc. We have also established the conditions for which the scalar curvature is harmonic, sub-harmonic and super-harmonic. We also established the existence condition of a gradient $\rho$-Einstein soliton in the Riemannian 3-manifolds, and consequently we proved some results. To verify our results, we constructed a non-trivial example of a three-dimensional Riemannian manifold equipped with a semi-symmetric metric $\xi$-connection. These topics are modern and have a lot of scope for researchers.

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